CHARM: Cassini-Huygens Mission to Saturn 10th Anniversary!!

Titan Highlights

Zibi Turtle, JHU/APL

26 August 2014

CHARM: Cassini's 10th Anniversary -- Titan!

7 Oct. 2013

	Four-Ye		ass Tour, Equi	ini I	Mis		n C	Vel		Septembe	r 2017		
Year of Tour	Pri 1 '04-'05	m e 2 '05-'06	Miss 3 '06-'07	i o n 4 '07-'08	Equino 5 '08-'09	x Mission 6 '09-'10	s 7 '10-'11	o I s 8 '11-'12	t i c 9 '12-'13	e 10 '13-'14	M i s 11 '14-'15	s i 12 '15-'16	<mark>o n</mark> 13 '16-'17
Orbits	11	15	22	27	39	21	16	19	25	12	12	20	56
Titan *Huygens							•••	•••					Proximal Orbits
Enceladus		•		•	9 9 9	•• ••	••	9 9 9 9 9 9				•• •	9 0 0 0 0
Other Icy Satellites (under 10,000 km)	W Phoebe	 Tethys Hyperion Dione Telesto Rhea 		 Rhea Iapetus Epimetheus 		 Mimas Rhea Helene Dione JG arc 	●Rhea ●Helene	 Dione Dione Tethys Methone Telesto 	Rhea		Dione Tethys	● Dione ♥ Epimetheus ノG arc	EOM Sep 15, 2017
Saturn (seen from Sun) 26 August 2014				CHARM	> : Cassini	's 10th Ar	niversar	v Titan	>($\mathbf{)}$	



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10 years of exploring Titan with Cassini

- Long-term, dedicated, complementary observations essential to understanding such a complex system...
- …to build up global views…
 - North pole in winter darkness when Cassini arrived
 RADAR SAR swaths now cover >40%
 Global ISS, VIMS, RADAR radiom./scatterometry
- ...& witness temporal variability (~1/3 Titan yr)
 Changes in weather patterns
 Changes in haze distribution (altitude, hemisphere)

Titan's Atmospheric Variability

detached haze 150 km higher than observed by *Voyager* Seasonal changes in weather patterns mid-latitude streaks

South-polar convective clouds

complex, variable haze structure

Titan's Atmospheric Variability

detached haze 150 km higher than observed

by Voyager

Returned to Voyager level around equinox Seasonal mid-latitude streaks changes in weather patterns Storms have yet to pick up at northern latitudes South-polar convective clouds

complex, variable haze structure

Titan's intricate albedo patterns reveal an Earth-like combination of surface processes



possible channels but no direct evidence of surface liquids



Titan's intricate albedo patterns reveal an Earth-like combination of surface processes

Surface liquids confirmed at high latitudes in July 2006 possible channels but no direct evidence of surface liquids

few impact craters

Still only 8 named impact structures (40-400 km); ~50 candidates

RADAR

VIMS

ISS

aeolian patterns

 Vast expanses of dunes at low latitudes

 Tectonic and cryovolcanic features identified

ectonic patterns

CIRS identifies propylene in Titan's lower atmosphere

ISS, 3 Nov 2013



Nixon et al. (2013):

 Thermal infrared spectra of Titan's stratosphere reveal propylene (or propene), the chemical that forms the plastic polypropylene (recycling code #5), at 100-200 km altitude



VIMS and ISS capture best near-IR views of Titan's lake district and seas

- VIMS near-infrared spectral imaging (1-5 µm) of Titan's north polar region (false-color)
- Lakes and seas are liquid methane and ethane
- Areas that appear orange might be evaporite deposits, Titan's organic equivalent of salt flats on Earth



VIMS, 12 Sept. 2013

26 August 2014

VIMS and ISS capture best near-IR views of Titan's lake district and seas



RADAR map of Titan's north pole

The Cassini RADAR images in this map were obtained in multiple operating modes with resolutions of 0.3-1.5 km, 2-10 km, and 40-200 km. False coloring is used to distinguish bodies of liquid hydrocarbon (blue-black) from dry land (brews) and does not represent the visual appearance of Titan's surface.

CHARM: Cassini's 10

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Seas Mapper

Towada Laci

6 August 201

Ligeia

Mare

ATTRACTO

Links to RADAR map & flyover movie

- http://photojournal.jpl.nasa.gov/catalog/PIA17655
- http://photojournal.jpl.nasa.gov/catalog/PIA17656
- http://saturn.jpl.nasa.gov/video/videodetails/?videoID=271
- http://photojournal.jpl.nasa.gov/archive/PIA17656anno-640.mov
- http://photojournal.jpl.nasa.gov/archive/PIA17656-640.mov

Continuing development of south polar cloud as southern winter progresses

ISS, 13 July 2013

West et al. (2013): Formed late 2011 to early 2012 ~300 km altitude



Cassini gravity and topography data constrain Titan's ice shell and interior ocean

Mitri et al. (2014):

- Dense, likely salty, ocean
- Rigid ice shell
 ~50-100 km thick
- Topographic variations due to tectonic structures and freezing at base
- Any methane outgassing must be localized

Clouds over Ligeia Mare





- Wind speeds ~3-4.5 m/s (7-10 mph)
- Start of northern summer storms??
- http://photojournal.jpl.nasa.gov/catalog/PIA18420
- http://saturn.jpl.nasa.gov/news/newsreleases/newsrelease2014081

26 August 2014

+44 hr

+12 hr

+32 hr

+39 hr

Seasonal changes - waiting for N summer storms...

Event, Date	<u>Time in Titan's year</u>
Voyager 1 flyby, Nov. 1980	29 March
Voyager 2 Flyby, Aug. 1981	8 April
Cassini SOI, 2 July 2004	Mid-January
Dissipation of high-altitude north-polar ethane cloud (VIMS), 2008-2009	Late N. winter
11 Aug 2009	N. vernal equinox
Decrease in altitude of detached haze (ISS), 2009-2010	Early N. spring
Low-latitude storm (ISS), SeptOct. 2010	Early April
Rapid changes in south polar upper atmospheric temperatures and composition (CIRS), 2010-2011	N. spring
South polar vortex (ISS, VIMS), 2011/2012	Late April
26 August 2014	Mid-May
Development of northern clouds (model predictions)	Spring? Summer??
May 2017	N. summer solstice



YEARS at SATURN

Huygens: first landing on a moon in the outer solar system

- Titan revealed as an Earth-like world with rain, rivers, lakes & seas; wind & dunes; tectonic, cryovolcanic, & impact structures
- Complex prebiotic chemistry and discovery of subsurface ocean

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