



Cassini MAPS highlights: Moon-magnetosphere interactions

Sarah Badman

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Magnetosphere

 The region which is controlled by the planet's magnetic field, containing a variety of charged particles (plasma).

Image: NASA

Moons interact with the magnetic field and plasma



Hyperion



- Vital statistics:
- size 180 x 133 x 103 km
- low density mainly water ice?
- orbits at ~25 R_S (1 R_S = Saturn radius = 60268 km)

Hyperion in Saturn's magnetic field



 Saturn's magnetic field rotates over Hyperion with a rotation period of ~11 h

Electron absorption



- Saturn's magnetic field lines move past Hyperion
- Electrons travelling along the magnetic field lines are absorbed by the moon

Surface charging



- Magnetic field lines move past Hyperion
- Electrons travelling along the magnetic field lines are absorbed by the moon
- The electrons make the moon surface negatively charged

Electrons repelled



- Magnetic field lines move past Hyperion
- Electrons travelling along the magnetic field lines are absorbed by the moon
- The electrons make the moon surface negatively charged
- The negatively charged surface repels other electrons, accelerating them away from the moon

Cassini measurements





When Cassini flew through the magnetic field lines connected to Hyperion it measured the repelled electrons

Cassini measurements





 Cassini also detected a drop-out in the higher energy electrons which had been absorbed by the moon

Hyperion: an electron accelerator



- Hyperion's surface was electrically charged to around -200 volts by plasma from Saturn's magnetosphere striking it.
- Low energy electrons were accelerated up to the spacecraft by the large potential difference. Cassini can remotely detect charging conditions on moons.
- Surface charging of a natural body has previously only been observed at the Earth's Moon first published detection in the Outer Solar System.
- Nordheim et al., GRL, 2014.

Titan



- Vital statistics
- diameter ~5000 km
- thick nitrogen atmosphere
- orbits at ~20 $R_{\rm S}$ (1 $R_{\rm S}$ = Saturn radius = 60268 km)

Magnetic field draping



 Saturn's magnetic field drapes around Titan's thick atmosphere

Effect of sunlight



 Sunlight ionises Titan's atmosphere ("photoionisation")

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Charge separation



- Sunlight ionises Titan's atmosphere ("photoionisation")
- Electrons (-) separate from ions (+) in the upper atmosphere
- The electrons are more energetic than the ions and can move along the magnetic field lines

Electron energy

- (a) energy ectron
- The electron energy is characteristic of the nitrogen atmosphere
- Cassini detects the electrons (photoelectrons) far away from where they are produced



Dayside and tail



Cassini detects the electrons from the atmosphere on many Titan flybys, both on through the sunlit region and in the distant tail:



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Titan's 'polar wind'



- The separation of the electrons

 (-) and ions (+) sets up an
 electric field, which pulls the
 ions (+) out of the atmosphere
- The voltage between the sunlit atmosphere and the tail is ~3 V
- Cassini measurements show Titan is gradually losing its atmosphere to the surrounding space at a rate of ~7 tonnes every 24 h

Coates et al., GRL, 2012; 2015. 20

Summary

- Saturn's moons interact with the planet even though they don't have their own magnetic fields
- The irregular satellite Hyperion has a charged surface, which accelerates electrons away from it
- Titan's atmosphere is ionised by sunlight, leading to loss of mass along Saturn's magnetic field lines draped through its atmosphere

Coming up

- Cassini will soon move into the last phase of its mission: the Grand Finale
- The last close flyby of Hyperion by Cassini took place on 31st May 2015
- There will be several more close Titan flybys until mid-2017 so we can study the moon's interaction over different seasons

References

- Nordheim, T. A., G. H. Jones, E. Roussos, J. S. Leisner, A. J. Coates, W. S. Kurth, K. K. Khurana, N. Krupp, M. K. Dougherty, and J. H. Waite (2014), Detection of a strongly negative surface potential at Saturn's moon Hyperion, Geophys. Res. Lett., 41, 7011–7018, doi:10.1002/2014GL061127.
- Coates, A. J., A. Wellbrock, J. H. Waite, and G. H. Jones (2015), A new upper limit to the field-aligned potential near Titan, Geophys. Res. Lett., 42, 4676–4684, doi:10.1002/2015GL064474.
- Coates, A. J., et al. (2012), Cassini in Titan's tail: CAPS observations of plasma escape, J. Geophys. Res., 117, A05324, doi:10.1029/2012JA017595.